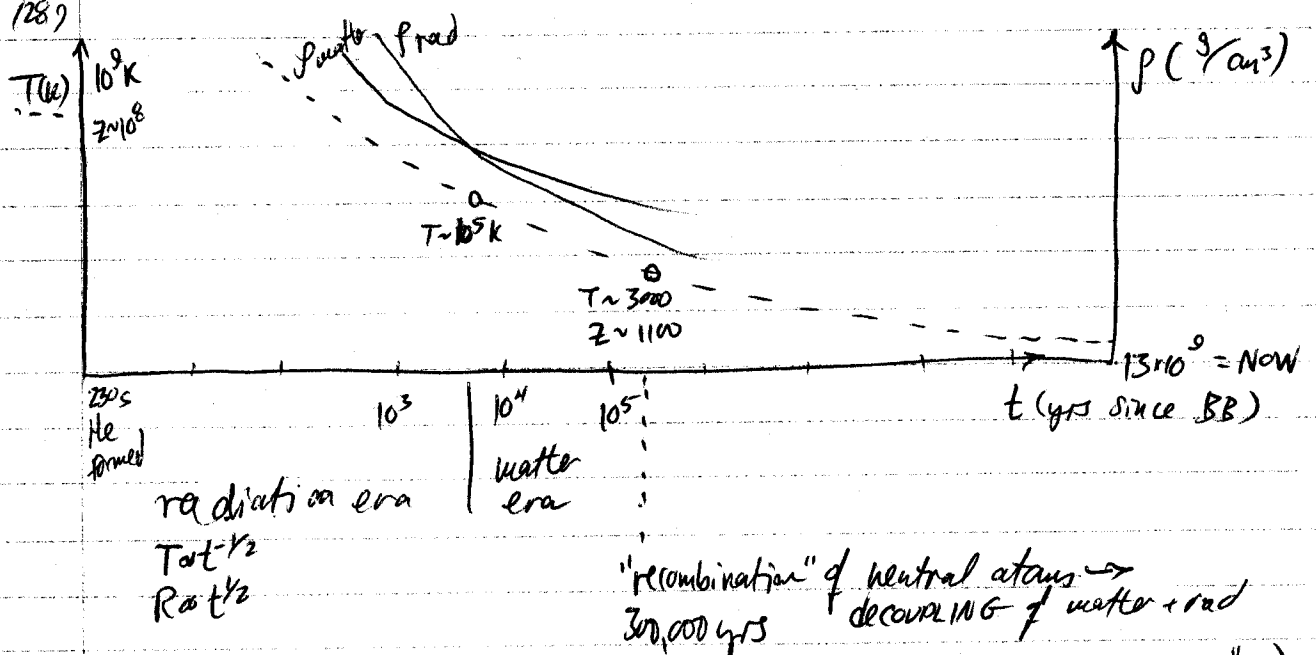


Ch 28 - THE EARLY UNIVERSE

§28.1 CONSTRUCTING THE UNIVERSE



B3FH: BIG BANG NUCLEOSYNTHESIS in RADIATION ERA (before 10" s)

He abundance due to fusion in first 3 minutes (not much from stars)

At $T \sim 10^{12} \text{ K}$, $\frac{n_n}{n_p} = 0.985$ (Boltzmann $\frac{28.4}{290}$)

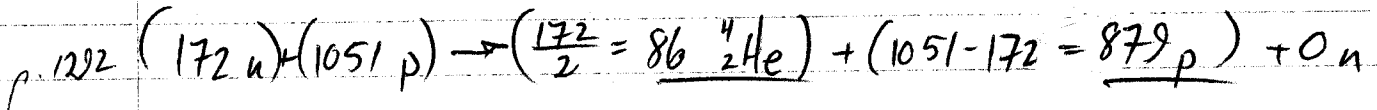
Around $T \sim 10^{10} \text{ K}$, $E_\gamma < 2m_e$: " $e^- + e^+$ annihilate, leaving too few e^- to drive $p + e^- + \bar{\nu}_e \rightarrow n$ " (plus ν_e drops)

$$\frac{n_n}{n_p} = \frac{223}{1000} \rightarrow \frac{172}{1051}$$

$n \rightarrow p + \bar{\nu}_e$

$10^{10} \text{ K} \qquad 10^9 \text{ K}$

p. 1291 Below 10^9 K , He nuclei form = MOST TIGHTLY BOUND light nucleus



$$\frac{M_{\text{He}}}{M_{\text{tot}}} = \frac{4 \cdot N_{\text{He}}}{N_p + 4 \cdot N_{\text{He}}} = \frac{4 \cdot 86}{879 + 4(86)} = 0.28$$

(PRE)

GALAXY + CLUSTER Structures form in MATTER ERA

p. 1293

ADIABATIC fluctuations

$$\Delta = \frac{\delta p}{p} \text{ same for RAD + MATTER}$$

ISOTHERMAL fluctuations

$$\Delta = \frac{\delta p}{p} = 0 \text{ for everything}$$

(constant density)

p. 1295

$$\Delta \propto t \text{ (RAD era)}, \Delta \propto t^{2/3} \text{ (Matter Era)}$$

preserves Δ from dissipation of acoustic osc.
no changes until recombination

p. 1296

JEAN'S MASS = M_{\min} for self-grav.

Before
decoupling:
 $v_s \sim \frac{c}{\sqrt{3}}$

$$M_J \propto \frac{\rho_{\text{baryons}} v_s^3}{\rho^{3/2}} \propto \frac{T^3}{(T^4)^{3/2}} \propto \frac{1}{T}$$

ρ_{photons}

until recombination: $M_J \propto T^{3/2}$

Δ can grow
After recombination:
 $M_J \sim 10^6 M_{\odot}$ (p. 1299)

p. 1297

RAD era: $\Delta = \text{const}$

MATTER era: ACOUSTIC OSCILLATIONS!

SIZE OF Globular Clusters
(stellar)

p. 1298

smooth out small $\Delta <$ photodistance

Surviving Δ have $M_{\min} \sim 10^{13} M_{\odot}$

SIZE of small GAL. CLUSTERS

p. 1299

Ex 1: Observed radio-quiet quasars at $z \sim 5$ require
 $\Delta \sim 10^{-3}$ at recombination (decoupling)

p. 1300

HOWEVER, anisotropies in CBR imply $\Delta_{\text{recom}} \sim 10^{-5}$

BARYONIC matter cannot clump fast enough to make observed structures
 \therefore 99% is DARK MATTER

SURPRISE! LOWER Δ CBR regions: COOLER (gas) $<$ HOTTER (grav redshift)!

p. 1302

DARK MATTER $\approx 94\%$

($\rho_{\text{baryonic}} < 6\% \rho_{\text{univ}}$)

HOT DARK MATTER

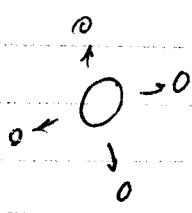
neutrinos : $m \sim 7-25 \text{ eV}$

$N_{\nu} \sim 10^{90} = \# \text{ of primordial photons}$

Ex: calculate $\Sigma m_{\nu} =$ _____

$\frac{\rho_{\nu}}{\rho_c} =$ _____

Like ADIABATIC Δ ,
 ν resist clumping \therefore
TOP DOWN formation



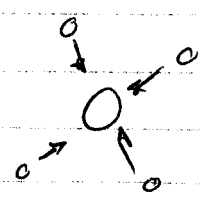
COLD DARK MATTER

WIMPS : $m \sim 10^{10-12} \text{ eV}$

AXIONS : $m \sim 10^{-5} \text{ eV}$

NEITHER HAS BEEN OBSERVED

COLD DM clumps more readily \therefore
BOTTOM-UP formation



perhaps a better fit to observations?

p. 1310 However, breakup into galaxies may occur too late

However, hard to account for voids

1319

SYMMETRY BREAKING \rightarrow COSMIC DISCONTINUITIES?

0D: monopole;
 10^{16} GeV

1D: cosmic string
 10^{21} g/cm

2D: domain wall

3D texture

14 Feb 82?

False vacuum moving at $v \sim c$?

INFLATION: See Ch 20 Weinberg lecture

28.13 In Problem 9.3, you derived Eq. (9.60) for the number density of black-body photons,

$$n = \frac{u}{2.70kT} = \frac{aT^3}{2.70k} = \frac{\text{photons}}{\text{volume}}$$

Use this result along with the baryon density, $\rho_{B,0}$, to estimate the ratio of the number of baryons to the number of photons in today's universe. For convenience, assume that the universe is made solely of hydrogen.

use $T = 2.73 \text{ K}$

28.14 Estimate the thickness of a typical cosmic string.

(2.5)
258

$$u = \frac{4\sigma T^4}{c} = aT^4$$

(9.60)
307

$$\frac{u}{n} = \frac{\pi^4 kT}{15(2.70k)} = 2.70 kT$$

From Prob 9.3:

$$n_{\text{photons}} =$$

present $n_{\text{baryons}} = n_{B,0} = \frac{\rho_{B,0}}{m_H} \leftarrow (27.16) \rho_{\text{crit}}$
 $m_H \leftarrow$ assume hydrogen universe

$$n_{B,0} =$$

$$\frac{u_{B,0}}{u_{\text{photons}}} =$$

28.14 Cosmic strings have density $\approx \lambda \sim 10^{21} \text{ g/cm} = \frac{\text{mass}}{\text{vol. length}}$

False vacuum has energy density = $\frac{\text{energy}}{\text{vol}} \approx 10^{25} \frac{\text{erg}}{\text{cm}^3}$ (28.13)
1312

$$\frac{\text{Energy}}{\text{length}} = \frac{mc^2}{\text{length}} =$$

Assume Area = πr^2 and solve for r :